

*Open topics in applied mathematics:*  
*Fast Methods in Scientific Computation*  
*APPM4720 / 5720*

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(These notes will be posted on the class webpage.)

## Purpose of class:

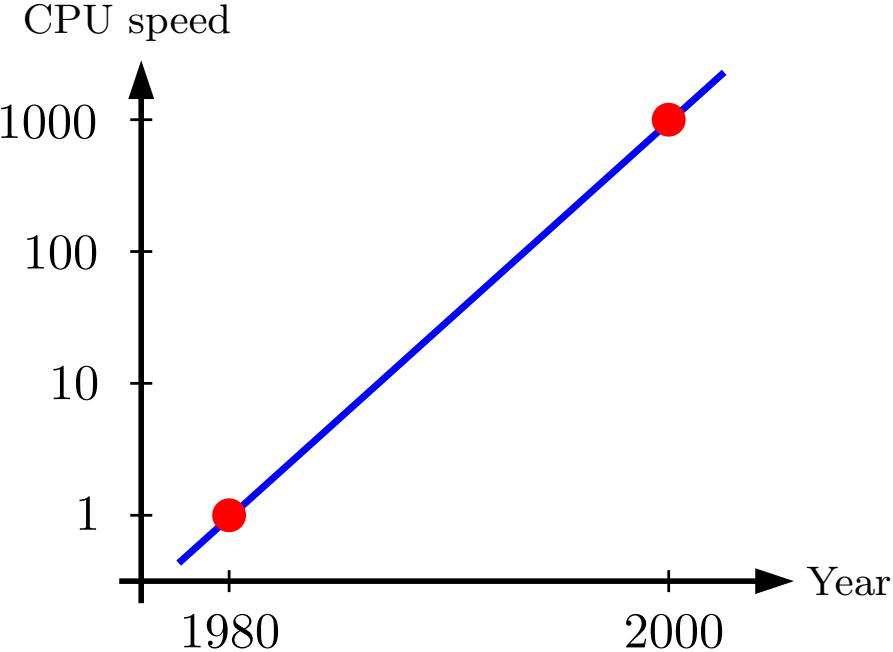
- The central theme of the class is to describe “fast” methods for solving elliptic PDEs such as
  - The Laplace and Poisson equations.
  - Helmholtz’ equation.
  - Maxwell’s equation.
  - The equations of linear elasticity.
  - The Navier-Stokes, and the Stokes equations.
- The class will also discuss several computational methods with broader use:
  - Quadrature. FFT. Fast methods for  $N$ -body problems.
- Practical programming techniques will be important. A principal goal is to encourage computing: How do you minimize the pain and maximize the fun?
- Math, modeling, and numerical methods will all be emphasized.

## Definition of the term “fast”:

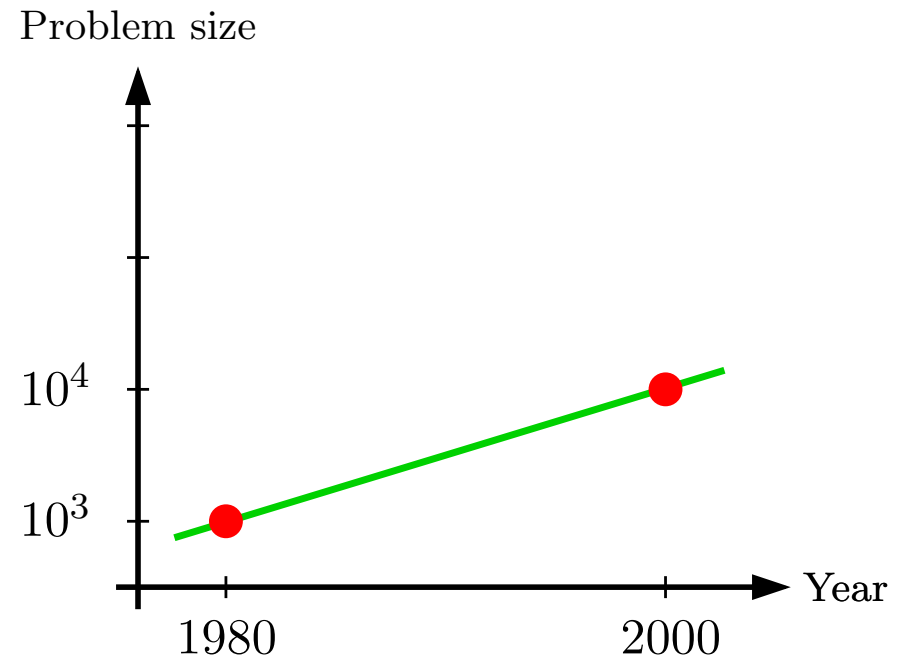
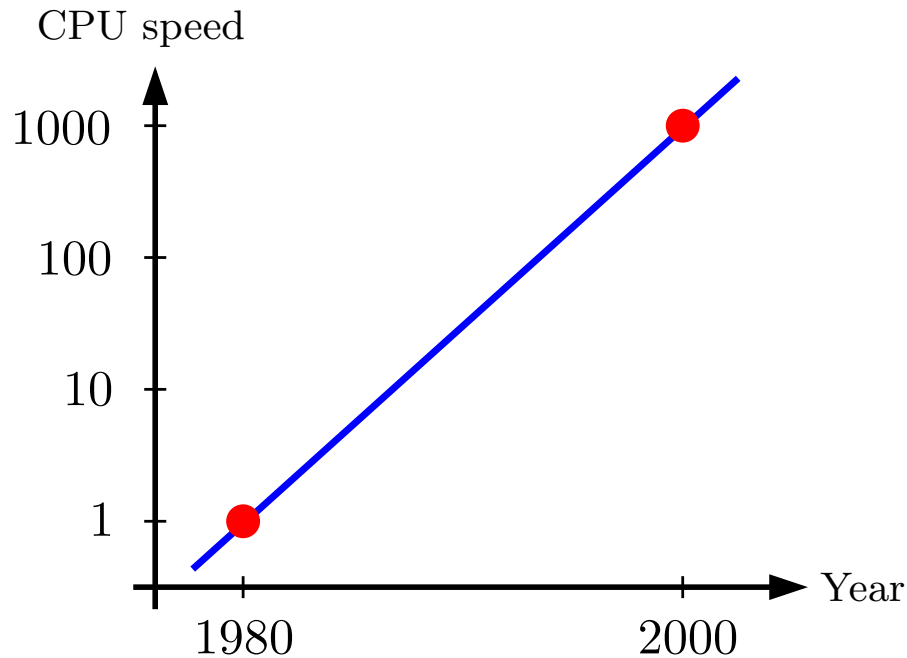
We say that a numerical method is *fast* if its execution time scales as  $O(N)$  as the problem size  $N$  grows.

Methods whose complexity is  $O(N \log N)$  or  $O(N \log^2 N)$  are also called “fast”.

# Growth of computing power and the importance of algorithms



## Growth of computing power and the importance of algorithms

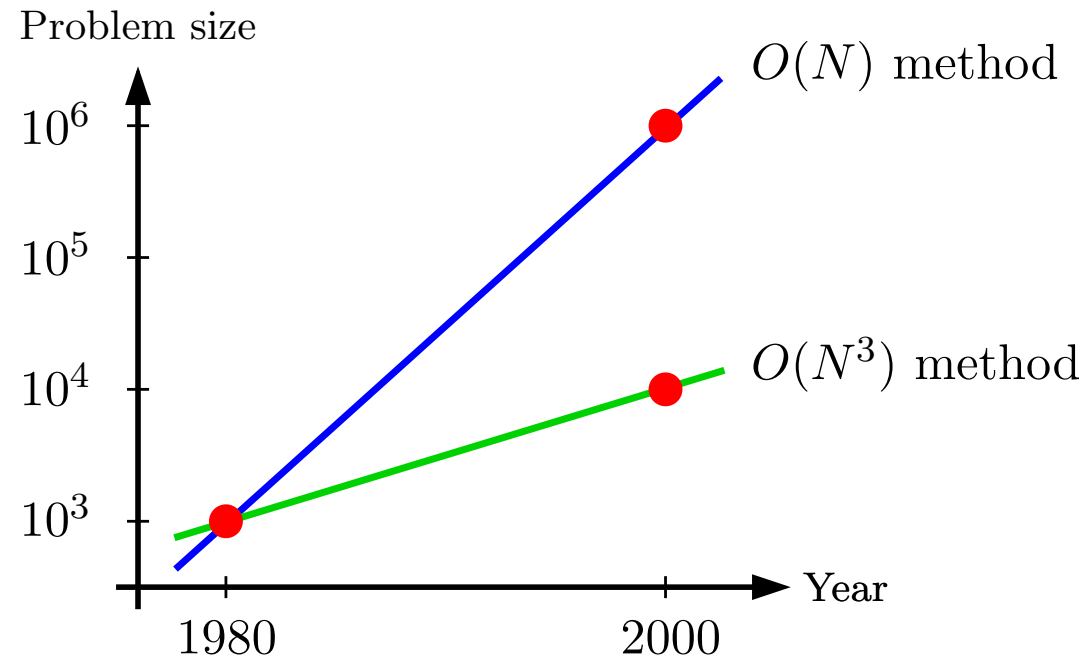
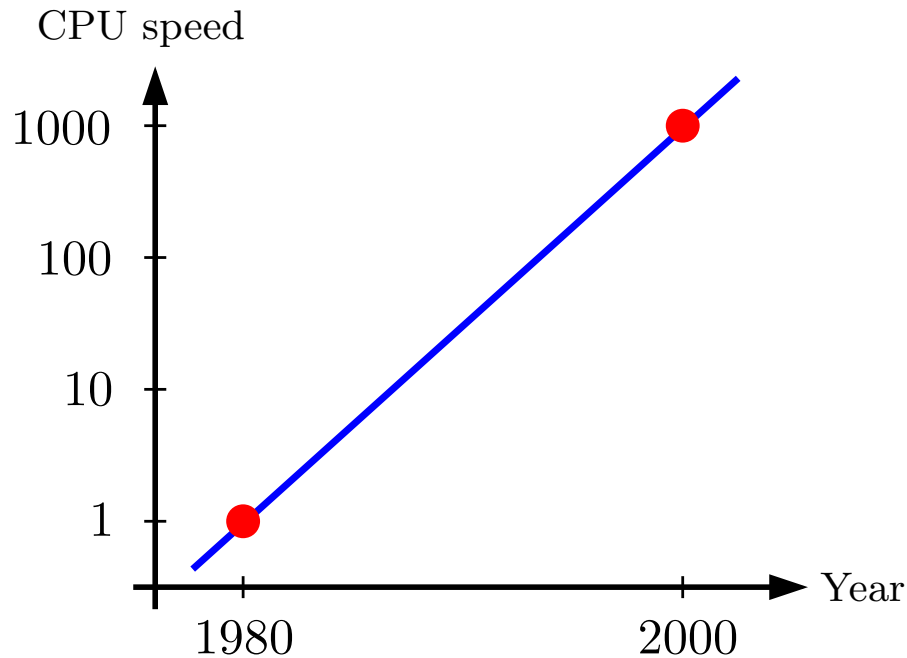


Consider the computational task of solving a linear system  $A\mathbf{u} = \mathbf{b}$  of  $N$  algebraic equations with  $N$  unknowns.

Classical methods such as Gaussian elimination require  $O(N^3)$  operations.

Using an  $O(N^3)$  method, an increase in computing power by a factor of 1000 enables the solution of problems that are  $(1000)^{1/3} = 10$  times larger.

## Growth of computing power and the importance of algorithms



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*Using a method that scales as  $O(N)$ , problems that are 1000 times larger can be solved.*

**Caveat:** It appears that Moore's law is no longer operative.

Processor speed is currently increasing quite slowly.

The principal increase in computing power is coming from *parallelization*.

In consequence, successful algorithms must scale well both with problem size and with the number of processors that a computer has.

To slightly offset the difficulty of parallelization, the *cost of storage is decreasing*. However, the speed of access is increasing only slowly, again reinforcing the need to keep data local in designing algorithms.

## Laplace's equation (in two dimensions for simplicity)

Let  $u = u(\mathbf{x})$  denote a differentiable function of the vector valued variable  $\mathbf{x} = (x_1, x_2)$ . The *Laplace operator* “ $-\Delta$ ” is defined by

$$\Delta u = \frac{\partial^2 u}{\partial x_1^2} + \frac{\partial^2 u}{\partial x_2^2}.$$

Let  $\Omega$  denote a *domain* with boundary  $\Gamma$ . Then the *Poisson equation* on  $\Omega$  is

$$(1) \quad \begin{cases} -\Delta u(\mathbf{x}) = f(\mathbf{x}), & \mathbf{x} \in \Omega, \\ u(\mathbf{x}) = g(\mathbf{x}), & \mathbf{x} \in \Gamma. \end{cases}$$

The function  $f$  is a given *body load* and  $g$  is a given *boundary data*. If  $f = 0$ , we call (1) the *Laplace equation*.

The Poisson and Laplace equations are the simplest equations in a large class of so called *elliptic PDEs*. Other examples include Helmholtz, elasticity, Maxwell (for the “time-harmonic case”).



The *Laplace and Poisson* equations:

**Electrostatics:**

$$\begin{cases} -\Delta u(\mathbf{x}) = f(\mathbf{x}), & \mathbf{x} \in \Omega, \\ u(\mathbf{x}) = g(\mathbf{x}), & \mathbf{x} \in \Gamma. \end{cases}$$

$u$  is the electric potential

$f$  is the electric charge density

$g$  is a fixed potential on the boundary (Neumann b.c.  $\Rightarrow$  fixed fluxes)

Examples of applications:

- Design of electric engines / turbines / etc.
- Biochemical modeling.
- Design of electronic circuits.

(“Magnetostatics” is entirely analogous.)

The *Laplace and Poisson* equations:

**Gravity:**

$$-\Delta u(\mathbf{x}) = f(\mathbf{x}), \quad \mathbf{x} \in \mathbb{R}^3,$$

$u$  is the gravitational potential

$f$  is the mass density

Examples of applications:

- Astrophysics

*A “hidden” Laplace problem:* Consider a situation with  $N$  gravitational bodies in  $\mathbb{R}^3$ . Each body has mass  $m_i$  and location  $\mathbf{x}_i$ . Then the force on body  $i$  resulting from interactions with the other bodies is

$$\mathbf{F}_i = \sum_{j \neq i} G m_i m_j \frac{\mathbf{x}_i - \mathbf{x}_j}{|\mathbf{x}_i - \mathbf{x}_j|^3},$$

where  $G \approx 6.67428 \cdot 10^{-11} \text{m}^3/(\text{kg s}^2)$  is the gravitational constant.

We now observe that the force  $\mathbf{F}_i$  can be expressed as

$$\mathbf{F}_i = -m_i \sum_{j \neq i} \nabla u_j(\mathbf{x}_i),$$

where  $u_j = u_j(\mathbf{x})$  is the gravitational potential generated by the  $j$ 'th charge

$$u_j(\mathbf{x}) = \sum_{j \neq i} G m_j \frac{1}{|\mathbf{x} - \mathbf{x}_j|}.$$

The potential  $u_j$  satisfies

$$-\Delta u_j(\mathbf{x}) = m_j \delta(\mathbf{x} - \mathbf{x}_j).$$

The total field  $u = \sum_i u_i$  satisfies

$$-\Delta u(\mathbf{x}) = f(\mathbf{x}) = \sum_i m_j \delta(\mathbf{x} - \mathbf{x}_j).$$

The problem of computing a sum such as

$$\mathbf{F}_i = \sum_{j \neq i} G m_i m_j \frac{\mathbf{x}_i - \mathbf{x}_j}{|\mathbf{x}_i - \mathbf{x}_j|^3},$$

arises directly in many applications:

- Astrophysics.
- Biochemical simulations (each “particle” is a charged part of a molecule).
- Modeling of semi-conductors (each “particle” is an ion).
- Fluid dynamics (each “particle” is an “vortex”).

It also arises indirectly in many “fast” methods for solving elliptic PDEs.

The naïve computation of  $\{\mathbf{F}_i\}_{i=1}^N$  requires  $O(N^2)$  operations since there are  $N(N-1)/2$  “pair-wise interactions.”

We will study in some detail a method that requires only  $O(N)$  operations; the so called *Fast Multipole Method* or *FMM*.

The *Laplace and Poisson* equations:

**Thermostatistics:**

$$\begin{cases} -\Delta u(\mathbf{x}) = f(\mathbf{x}), & \mathbf{x} \in \Omega, \\ u(\mathbf{x}) = g(\mathbf{x}), & \mathbf{x} \in \Gamma. \end{cases}$$

$u$  is the temperature

$f$  is the heat source density

$g$  is a fixed temperature on the boundary (Neumann b.c.  $\Rightarrow$  fixed flows)

Examples of applications:

- ...

The *Helmholtz* equation:

Recall the *wave equation*:

$$\frac{\partial^2 u}{\partial x_1^2} + \frac{\partial^2 u}{\partial x_2^2} = \frac{\partial^2 u}{\partial t^2}.$$

The wave equation models vibrations in membranes, acoustic waves, certain electro-magnetic waves, and many other phenomena.

Now assume that the time dependence is “time harmonic”:

$$u(\mathbf{x}, t) = v(\mathbf{x}) \cos(\omega t).$$

Then  $\frac{\partial^2 u}{\partial t^2} = -\omega^2 u$  and so the wave equation becomes the *Helmholtz equation*:

$$\frac{\partial^2 v}{\partial x_1^2} + \frac{\partial^2 v}{\partial x_2^2} = -\omega^2 v.$$

The *Maxwell equations*

$$\begin{cases} \nabla \cdot \mathbf{E} = \rho & \nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t} \\ \nabla \cdot \mathbf{B} = 0 & \nabla \times \mathbf{B} = \mathbf{J} + \frac{\partial \mathbf{E}}{\partial t} \end{cases}$$

model electromagnetism.  $\mathbf{E}$  is the electric field, and  $\mathbf{B}$  is the magnetic field.

Consider the stationary case where  $\partial \mathbf{E} / \partial t = 0$  and  $\partial \mathbf{B} / \partial t = 0$ . Since  $\mathbf{E}$  is curl-free, there exists a function  $u = u(\mathbf{x})$  such that

$$\mathbf{E} = -\nabla u.$$

(The function  $u$  is the electric potential.) We now find that

$$\rho = \nabla \cdot \mathbf{E} = \nabla \cdot (-\nabla u) = -\Delta u,$$

and we recover the Poisson equation we saw earlier:

$$-\Delta u = \rho.$$

The *Maxwell equations*

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Now consider another simplification: the “time-harmonic” case where

$$\mathbf{E}(\mathbf{x}, t) = \mathbf{E}(\mathbf{x}) e^{i\omega t}, \quad \mathbf{B}(\mathbf{x}, t) = \mathbf{B}(\mathbf{x}) e^{i\omega t}.$$

Then

$$\frac{\partial}{\partial t} \mathbf{E} = i\omega \mathbf{E}, \quad \text{and} \quad \frac{\partial}{\partial t} \mathbf{B} = i\omega \mathbf{B}.$$

Inserting these relations into the Maxwell equations, we obtain a system of “Helmholtz-like” equations

$$\begin{cases} \nabla \cdot \mathbf{E} = \rho & \nabla \times \mathbf{E} = -i\omega \mathbf{B} \\ \nabla \cdot \mathbf{B} = 0 & \nabla \times \mathbf{B} = \mathbf{J} + i\omega \mathbf{E} \end{cases}$$

(In special cases, the system simplifies to the plain Helmholtz equation.)



Recall:

$$\begin{cases} \nabla \cdot \mathbf{E} = \rho & \nabla \times \mathbf{E} = -i\omega\mathbf{B} \\ \nabla \cdot \mathbf{B} = 0 & \nabla \times \mathbf{B} = \mathbf{J} + i\omega\mathbf{E} \end{cases}$$

Suppose  $\rho = 0$  and  $\mathbf{J} = \mathbf{0}$ . Then

$$\nabla \times \nabla \times \mathbf{E} = \nabla \times (-i\omega\mathbf{B}) = -i\omega(\nabla \times \mathbf{B}) = -i\omega(i\omega\mathbf{E}) = \omega^2\mathbf{E}.$$

Now recall that for any vector field  $\mathbf{F}$  we have

$$\nabla \times \nabla \times \mathbf{F} = \nabla(\nabla \cdot \mathbf{F}) - \Delta\mathbf{F}.$$

Consequently:

$$\nabla(\nabla \cdot \mathbf{E}) - \Delta\mathbf{E} = \omega^2\mathbf{E}.$$

Finally recall that  $\nabla \cdot \mathbf{E} = 0$  to obtain the “Helmholtz-like” equation

$$-\Delta\mathbf{E} = \omega^2\mathbf{E}.$$

The equations of *linear elasticity* in  $\mathbb{R}^d$ :

$$\sum_{j,k,l=1}^d \frac{1}{2} E_{ijkl} \left( \frac{\partial^2 u_k}{\partial x_l \partial x_j} + \frac{\partial^2 u_l}{\partial x_k \partial x_j} \right) = f_i, \quad i = 1, 2, \dots, d.$$

The function  $\mathbf{u} = \mathbf{u}(\mathbf{x}) = (u_1(\mathbf{x}), u_2(\mathbf{x}), \dots, u_d(\mathbf{x}))$  is the displacement of an elastic material subjected to the body load  $\mathbf{f} = \mathbf{f}(\mathbf{x})$  at the point  $\mathbf{x}$ .

$(E_{ijkl})_{i,j,k,l=1}^d$  is the *stiffness tensor* which describes the material properties.

Many simplifications can be derived from the basic equilibrium equation. For instance, if the material is isotropic, and if  $\mathbf{f} = 0$ , then the displacements satisfy the *biharmonic equation*

$$(-\Delta)^2 \mathbf{u} = 0.$$

Another simplification is the displacement of a thin elastic membrane:

$$\begin{cases} (-\Delta)^2 u(\mathbf{x}) = f(\mathbf{x}), & \mathbf{x} \in \Omega, \\ u(\mathbf{x}) = g(\mathbf{x}), & \mathbf{x} \in \Gamma, \\ u_n(\mathbf{x}) = h(\mathbf{x}), & \mathbf{x} \in \Gamma. \end{cases}$$

Here  $f$  is the body load (e.g. gravity),  $h$  is the prescribed deflection at the boundary, and  $h$  is the prescribed normal derivative. (Since the equation has order *four*, we need *two* boundary conditions.)

## **Practicalities:**

### *Class notes:*

Notes will be posted on the class website. (There is no official text book.)

### *Class attendance:*

Strongly encouraged.

### *Computer programming:*

Matlab will be used. If you do not have access to a computer with Matlab, please contact the instructor.

### *Grading:*

Based on short weekly homeworks and two longer projects. No final exam.